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Number 25

19 November 2010

Reconnaissance Mineralogy of the Eocene Mole Hill Diatreme, Rockingham County, Virginia

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ISSN 1061-1878

Virginia Museum of Natural History
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ISSN 1061-1878

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ABSTRACT

The Mole Hill diatreme consists of picrobasalt with abundant megacrysts (0.5 mm to 2cm in maximum dimension) of clinopyroxene, Mg-Al-Fe spinel, and (less abundant) olivine. Cognate minerals include microphenocryst and groundmass plagioclase, olivine, clinopyroxene and Fe-Ti-(Cr) spinel. Clinopyroxene is the most abundant megacryst phase. Overall, the clinopyroxene in the megacryst cores is a high-Al, low-Cr augite with Mg# 78-88. Sieve textured rims approach groundmass clinopyroxene compositions. Olivine occurs as megacrysts and also as small (0.1- 0.5 mm) crystals of indeterminate origin. These may be phenocrysts, xenocrysts, or both. All non-groundmass olivine is zoned, becoming Fe- and Ca-rich (and approaching the composition of groundmass olivine) rimward. The most primitive olivine has Fo~90 and NiO as high as 0.75wt.%. More typically, olivine is Fo78-88 with NiO <0.5 wt.%. The megacryst/xenocryst olivine cores have higher Mg# and lower CaO than groundmass olivine. Megacrystic spinels are notably low in Cr, with Cr# <1, and variable Mg# ranging from 52-74. This variation appears to be continuous, despite the lack of zoning in individual spinel xenocrysts. Plagioclase occurs only as a microphenocryst phase, with uniform An75 cores and rims as sodic as An58. Cognate clinopyroxene (Mg#67-78) is enriched in Ca and Ti relative to the megacrysts. Groundmass olivine has low NiO and high (0.3-0.6 wt.%) CaO. Groundmass spinels have ulvospinel contents near 50%, initially rising with Mg# (in Cr-rich microphenocrysts) then dropping. Although the lack of context for the megacrysts precludes a definitive understanding of their origin, megacryst chemistry (especially the low-Cr spinels and the overall abundance of clinopyroxene) suggests a clinopyroxene-rich source in the upper (e.g. spinel zone) continental lithosphere. This source is likely similar to the Al-augite suite clinopyroxenites and wehrlites that occur as xenoliths and as intrusive veins in composite xenoliths from alkali basalt provinces worldwide. Cognate mineral (groundmass minerals and microphenocrysts) compositions are consistent with crystallization from a slightly evolved alkali basalt melt.

INTRODUCTION

A localized series of Eocene alkaline subvolcanic necks, pipes, and dikes (Eocene Subvolcanic Suite, ESS) outcrops in Highland County, Virginia and adjacent Pendleton County, West Virginia (Rader et al., 1986; Southworth et al., 1993). A single isolated volcanic plug, Mole Hill, outcrops in Rockingham County, Virginia 60 km east of the main field (Fig. 1).

The suite is intrusive into folded sedimentary rocks of Paleozoic age. The rocks of the ESS range in composition from picrobasalt to rhyolite. Basaltic rocks are nepheline-normative and plot in the field of alkaline within-plate basalt (Wood 1980; Southworth et al., 1993). Available bulk element and isotope chemistry is consistent with a mantle

origin for the suite with little evidence for crustal contamination aside from obvious xenoliths of high-level lithologies (e.g. Paleozoic limestone and sandstone) (Southworth et al., 1993; Tso et al., 2004; Tso and Surber, 2006). The more silicic rocks of the ESS appear to have been emplaced during explosive eruptions, as borne out by the presence of breccia pipes and peperites (Rader et al., 1986; Tso and Surber, 2006) and by the recognition of correlative volcanic ash deposits in the Eocene rocks of the North Carolina Coastal Plain (Harris and Fullagar, 1989). Several of the mafic plugs, including Mole Hill, have been interpreted as diatremes. Although the tectonic setting for the emplacement of these rocks is far from obvious, they seem to follow a NW-trending lineament associated with a NW trending fracture set. This may indicate emplacement along reactivated orogen-parallel faults. Southworth et al. (1993) suggest that this reactivation is related to extension accompanying plate reorganizations between 37 and 53Ma.

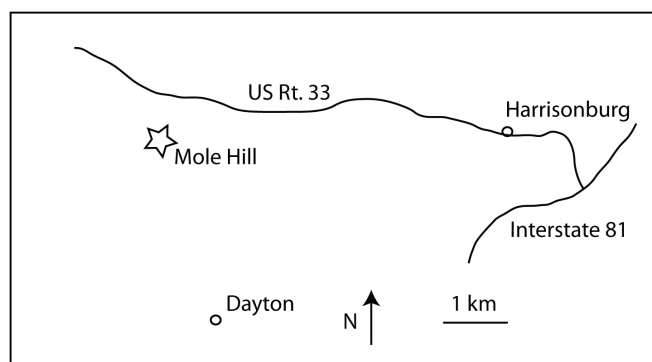


Figure 1: Location of Mole Hill in Rockingham County, Virginia.

The Mole Hill diatreme consists of alkali olivine microbasalt (tentative designation based on limited data) containing abundant large (0.5 mm to 2cm; large in the context of the host basalt) crystals of olivine, spinel and clinopyroxene (henceforth referred to as megacrysts) in a largely crystalline groundmass consisting of olivine, clinopyroxene, plagioclase, and Fe-Ti oxides. This study encompasses the description of the composition and compositional variability in the crystalline phases of the Mole Hill microbasalt including some interpretation of the origin of the megacryst phases.

Petrography and Mineral Chemistry

Methods. Minerals were analyzed using the Cameca SX-50 electron microprobe in the Department of Geological Sciences at Virginia Polytechnic Institute and State University. Beam current was 15na, at 20kv; spot size was 3-5 microns. Counting times were 20 seconds on peak and 10 seconds on two background points. A variety of natural and synthetic silicates and oxides were used as standards.

General. The basalt at Mole Hill is a alkali olivine microbasalt characterized by megacrysts of olivine, clinopyroxene, and aluminous spinel with a microphenocrysts of plagioclase, olivine, clinopyroxene and Fe-Ti-(Cr) oxides in a fine-grained holocrystalline groundmass of the same mineralogy. A weak to moderate alignment of plagioclase laths is apparent in some samples.

Olivine. Olivine occurs as megacrysts, a groundmass phase, and in crystals of intermediate size (0.1-0.5 mm) of, apparently, mixed xenocrystic/phenocrystic origin.

Megacryst/xenocryst cores range in (average) composition from Fo₈₂-Fo₉₀ (Table 1, Fig. 2). NiO content ranges from 0.2-0.6 wt.%, with some spot analyses in the most magnesian megacryst >0.7wt.% (Fig. 2). CaO concentrations in the megacrysts/xenocryst cores are <0.3 wt.% (Fig. 2). The megacrysts are zoned, with rims approaching groundmass olivine compositions and intermediate areas that may have partially equilibrated with the

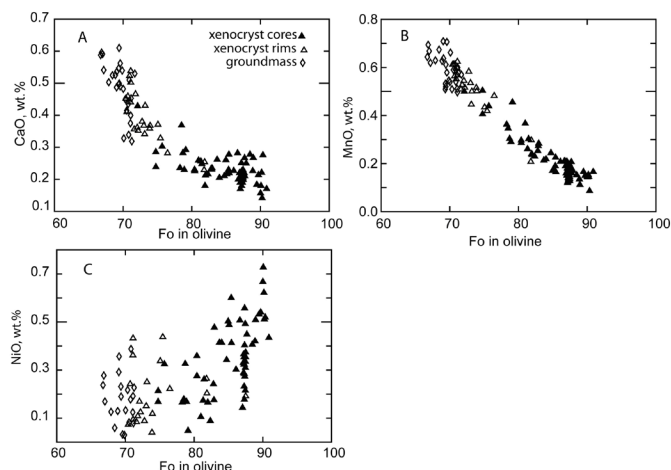


Figure 2. Variations in olivine chemistry, Mole Hill megacrysts and groundmass.

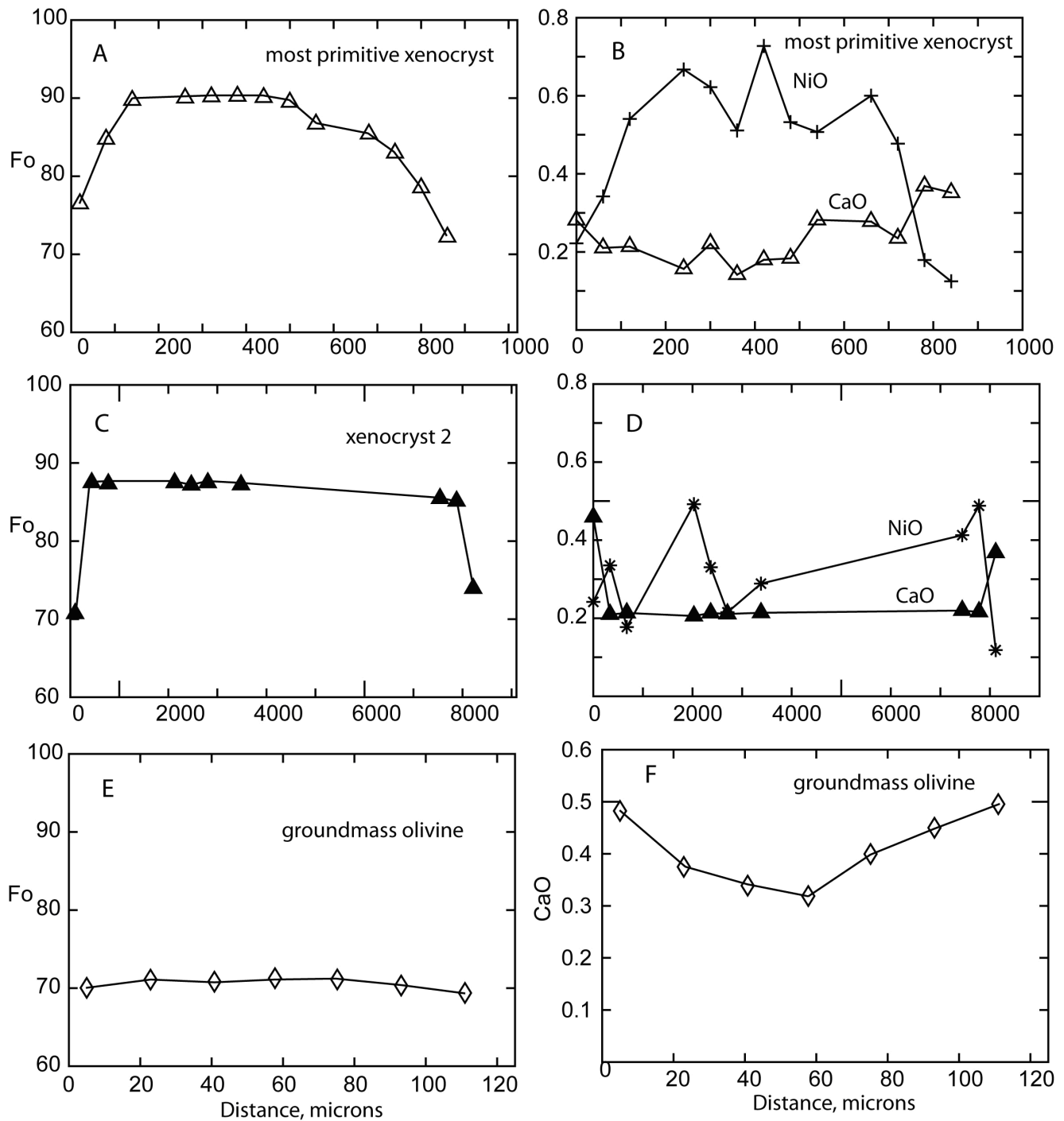


Figure 3. Chemical variation with distance along microprobe traverses in olivine. A, B. Most magnesian xenocryst. C, D. Large (1 cm) xenocryst. E, F. Microphenocryst.

Table 1. Olivine

VMNH #	81554	81564	81564	81554	81554	81564	81564	81564	81554	81554
sample	Mh 1-5-1	Mh 11	Mh 11	Mh 1-5-2	Mh 1-5-2	Mh 11	Mh 11	Mh 11	Mh 1-5-1	Mh 1-5-2
N	6	16	1	5	6	3	12	4	4	5
X'l type	mega	mega	Mega	mega	zoned	mega	phen	gmass	gmass	gmass
notes	core	v. large		core	core	rim				
SiO ₂	40.49	41.31	41.04	42.06	40.35	38.45	38.32	37.89	36.98	37.58
MgO	45.60	47.63	47.21	49.52	43.41	37.50	35.76	33.64	34.08	34.23
CaO	0.25	0.21	0.25	0.18	0.25	0.40	0.44	0.58	0.48	0.51
FeO	13.01	12.29	11.56	9.87	16.97	24.41	26.81	28.70	26.72	28.04
MnO	0.20	0.16	0.17	0.15	0.27	0.50	0.56	0.64	0.65	0.65
NiO	0.39	0.31	na	0.60	0.25	0.23	0.18	0.18	0.11	0.01
sum	99.93	101.92	100.23	102.38	101.50	101.49	102.07	101.62	99.01	101.02
Cations/4 oxygens										
Si	1.007	1.003	1.008	1.006	1.006	0.996	0.998	1.001	0.997	0.996
Mg	1.691	1.725	1.728	1.766	1.614	1.448	1.389	1.325	1.370	1.352
Ca	0.007	0.006	0.007	0.005	0.007	0.011	0.012	0.016	0.014	0.015
Fe	0.271	0.250	0.237	0.197	0.354	0.529	0.584	0.634	0.602	0.622
Mn	0.004	0.003	0.004	0.003	0.006	0.011	0.012	0.014	0.015	0.015
Ni	0.008	0.006	na	0.012	0.005	0.005	0.004	0.004	0.002	0.000
total	2.988	2.992	2.984	2.990	2.992	3.001	2.999	2.996	3.000	3.000
Fo	86.2	87.4	87.9	89.9	82.0	73.2	70.4	67.6	69.5	68.5

host basalt (e.g. Costa and Dungan, 2005) (Fig. 3). The low CaO and high NiO concentrations in the megacryst cores suggest that the megacrystic olivines are mantle xenocrysts that did not form in equilibrium with the host basalt (Simkin and Smith, 1970; Hirano et al., 2004; Rohrbach et al., 2005). Groundmass and (micro)phenocrystic olivine is substantially more Fe-rich (Fo_{68-80}) CaO-rich (CaO = 0.3-0.6 wt.%) and NiO-poor (NiO concentrations are usually below 0.3 wt. % with many values approaching detection limits (0.1 wt.%) than the megacryst cores (Table 1; Fig. 2). Zoning in true microphenocrysts is limited (i.e. Fo_{71-68} ; Fig. 3). More strongly zoned intermediate size and composition crystals (e.g. Fo_{80-70}) may either reflect an early phenocryst phase or partially re-equilibrated xenocrysts.

Clinopyroxene. Clinopyroxene is the most abundant megacryst and groundmass phase in the Mole Hill picobasalt. There are four chemically distinct populations of clinopyroxene in the Mole Hill picobasalt, with two of these represented by a single analyzed crystal. Most clinopyroxene megacrysts are high-Al (5.6-8.8 wt.% Al_2O_3) augites with TiO_2 <1.0 wt.% and Na_2O between 0.4-0.9 wt.% (Table 2; Fig. 4). Mg#'s of megacryst cores (average) range from 78-86. The most magnesian and Cr-rich clinopyroxene occurs in a two-crystal

“xenolith” with a large olivine xenocryst (Figure 4). The other analyzed xenocrysts vary in Mg# (Figs. 4,5), and can be characterized as low-Cr (e.g. Cr_2O_3 <0.2 wt.%). Most xenocrysts are zoned, with the sieve-textured rims approaching groundmass clinopyroxene composition. Groundmass clinopyroxene has Mg# near 70. It can be described as salite (Ca-rich augite) and is higher in Ti and, for the most part, lower in Al than the megacrysts. Some groundmass clinopyroxenes (and megacryst rims) have fairly high (>0.3 wt.%) Cr_2O_3 (Table 2; Fig. 4).

Two unusual clinopyroxene xenocrysts were analyzed. One has the low Mg# typical of groundmass clinopyroxene, but is distinctly lower in Ti and enriched in Na and Si with respect to the groundmass clinopyroxene (Table 2, Fig. 4). The other unique xenocryst is exceptionally enriched in Ti, Al and Ca, containing among the highest concentrations of these elements ever reported (Robinson, 1980) (Table 2). Similar high-Al pyroxenes have been reported as phenocrysts in alkaline igneous rocks (e.g. Gerke et al., 2005). This grain has a sieve-textured rim, suggesting reaction with the picobasalt host, and a xenocrystic (or, possibly, high-P phenocrystic) origin seems likely.

Spinel. Aluminous spinels (85-95% aluminous (i.e. sp + hc) end-member) occur as large (up to 5

Table 2: Clinopyroxene

VMNH #	81554	81554	81554	81554	81554	81554	81554	81564	81564	81564	81554	81554	81554	81564
sample	Mh 1-5-1	Mh 1-5-2	Mh lcp	Mh lcp	Mh lcp	Mh lcp	Mh lcp	Mh 11	Mh 11	Mh 11	Mh 1-5-1	Mh 1-5-2	Mh lcp	Mh 11
n	9	7	7	11	1	4	12	10	5	11	4	5	3	5
x'l type	mega	mega	mega	mega	mega	Xeno	*xeno cpx	mega	mega	mega	mega	mega	grdmass	grdmass
notes			core	intermed.	Rim	low-Al	high-Al	in ol xeno	in ol xeno	mega (rim)	grdmass	grdmass	grdmass	grdmass
SiO2	49.90	50.22	48.93	48.75	48.25	50.58	40.19	50.28	49.04	48.76	48.63	45.70	47.50	48.26
Al2O3	7.72	6.17	8.38	8.47	5.01	3.24	14.00	5.96	5.57	8.78	4.05	7.01	5.58	5.30
Na2O	0.65	0.53	0.90	0.82	0.50	1.03	0.24	0.53	0.43	0.78	0.47	0.65	0.59	0.47
MgO	16.53	15.43	15.17	14.68	13.20	12.58	9.82	16.54	14.61	15.02	12.65	11.61	12.50	12.66
CaO	18.68	21.65	18.42	19.45	23.44	21.54	24.74	20.94	22.52	19.69	22.43	22.29	22.98	22.92
TiO2	0.48	0.84	0.71	0.81	1.47	0.38	4.20	0.46	1.02	0.76	1.71	2.90	1.79	1.95
FeO	5.48	5.96	7.42	7.26	7.67	10.51	7.02	4.69	6.21	6.55	9.32	9.77	8.57	8.73
MnO	0.11	0.15	0.16	0.16	0.10	0.15	0.11	0.11	0.14	0.15	0.24	0.21	0.21	0.21
Cr2O3	0.18	0.04	0.03	0.02	0.00	0.07	0.00	0.61	0.47	0.04	0.01	0.10	0.14	0.14
sum	99.74	100.97	100.12	100.41	99.66	100.09	100.30	100.12	100.04	100.54	99.52	100.25	99.87	100.64
Cations/6 oxygens														
Si	1.822	1.832	1.797	1.790	1.817	1.905	1.519	1.838	1.820	1.783	1.842	1.731	1.793	1.805
Al	0.332	0.265	0.363	0.366	0.222	0.144	0.624	0.257	0.244	0.378	0.181	0.313	0.248	0.234
Na	0.046	0.037	0.064	0.059	0.037	0.075	0.017	0.038	0.031	0.055	0.035	0.048	0.043	0.034
Mg	0.900	0.839	0.831	0.804	0.741	0.707	0.553	0.902	0.808	0.819	0.714	0.656	0.704	0.706
Ca	0.731	0.846	0.725	0.765	0.946	0.869	1.002	0.820	0.896	0.772	0.911	0.904	0.930	0.919
Ti	0.013	0.023	0.020	0.022	0.042	0.011	0.119	0.013	0.029	0.021	0.049	0.083	0.051	0.055
Fe	0.167	0.182	0.228	0.223	0.241	0.331	0.222	0.143	0.193	0.200	0.296	0.309	0.271	0.273
Mn	0.004	0.005	0.005	0.005	0.003	0.005	0.003	0.004	0.004	0.005	0.008	0.007	0.007	0.007
Cr	0.005	0.001	0.001	0.001	0.000	0.002	0.000	0.017	0.014	0.001	0.000	0.003	0.004	0.004
total	4.020	4.030	4.033	4.034	4.049	4.049	4.059	4.031	4.038	4.034	4.036	4.053	4.051	4.038
AlIV	0.178	0.168	0.203	0.210	0.183	0.095	0.481	0.162	0.180	0.217	0.158	0.269	0.207	0.195
AlVI	0.154	0.097	0.160	0.156	0.039	0.049	0.143	0.095	0.063	0.161	0.023	0.043	0.042	0.039
Wo	40.6	45.3	40.6	42.7	49.0	45.6	56.4	44.0	47.2	43.1	47.4	48.4	48.8	48.4
En	50.0	45.0	46.6	44.9	38.4	37.1	31.2	48.3	42.6	45.8	37.2	35.0	37.0	37.2
Fs	9.3	9.7	12.8	12.4	12.5	17.4	12.5	7.7	10.2	11.2	15.4	16.6	14.2	14.4
Mg#	84.3	82.2	78.5	78.3	75.4	68.1	71.5	86.3	80.6	80.3	70.7	67.9	72.2	72.1

*High calculated Wo reflects high Al rather than high Ca

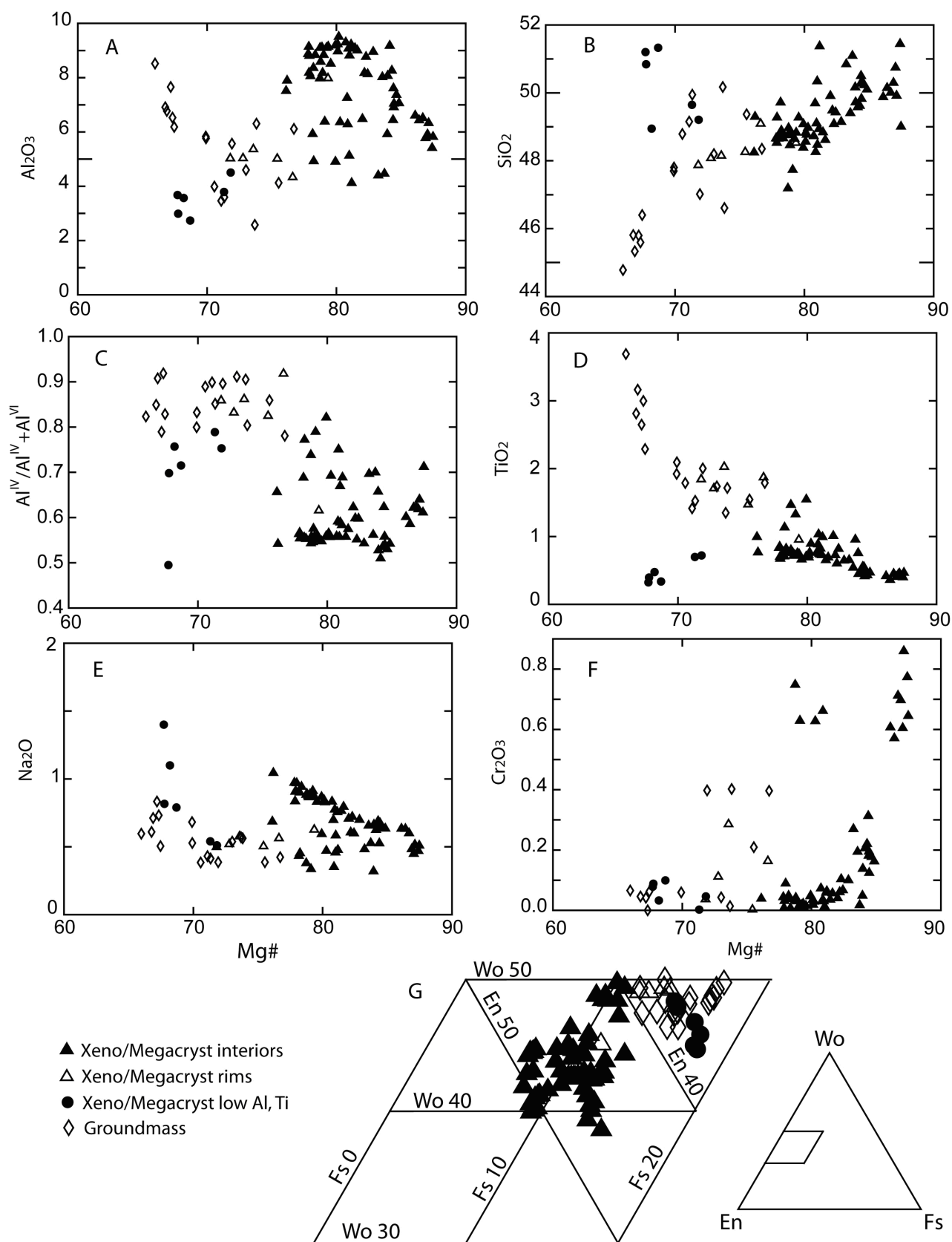


Figure 4. A-F Clinopyroxene chemical variation with Mg#. G. Plot of pyroxene quadrilateral compositions (uncorrected).

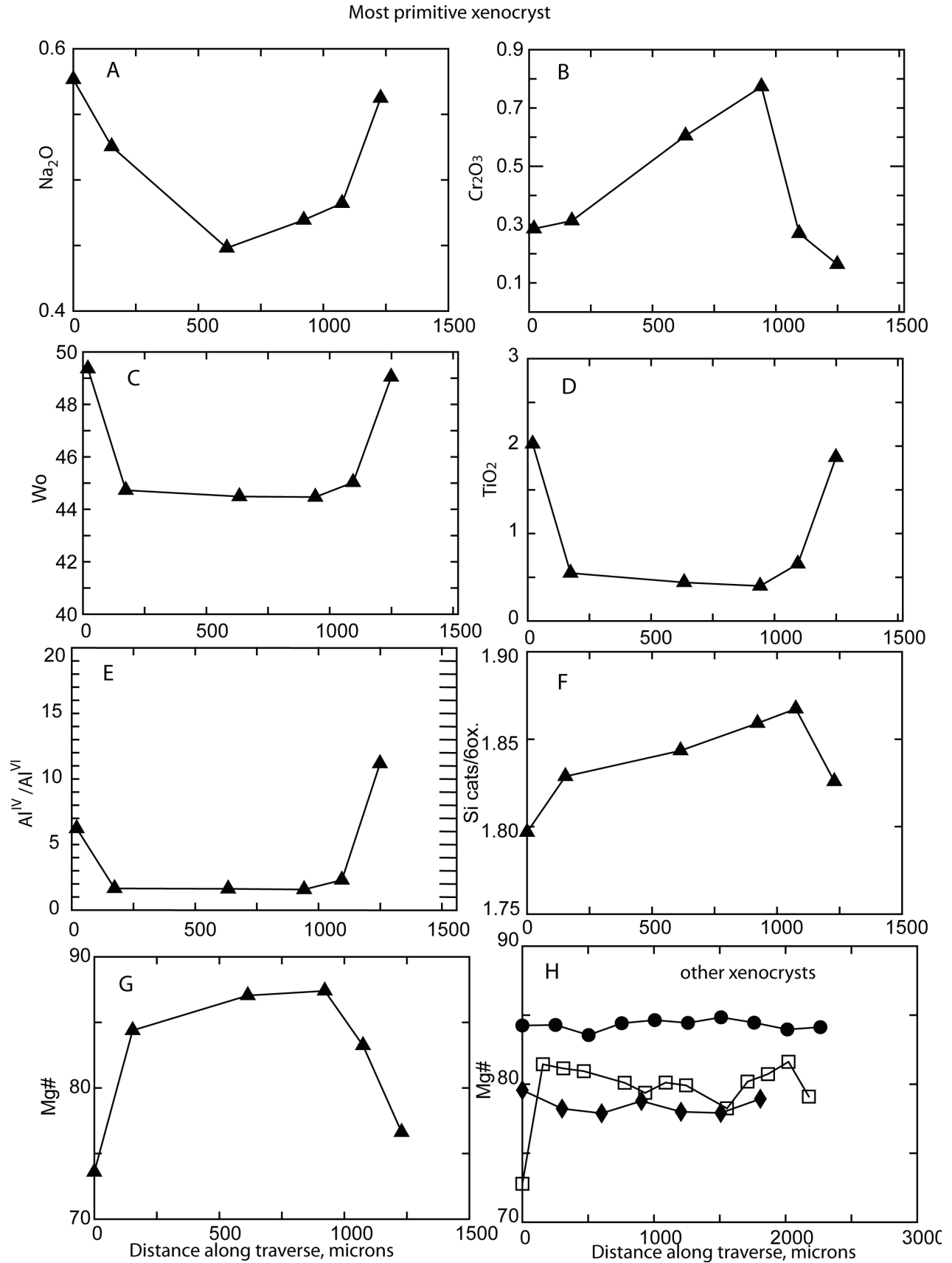


Figure 5. Chemical zoning along microprobe traverses. A-G, most Mg- and Cr-rich Clinopyroxene. H. Other clinopyroxene xenocrysts.

mm) rounded, pale greenish gray to nearly opaque crystals, with opacity correlating with total Fe. Most aluminous spinels are rounded, suggesting disequilibrium. The most Fe-rich aluminous spinels (e.g. Table 3, column 1) tend to be less rounded or even subhedral, but are also embayed. Aside from the rounding or embayment, the aluminous spinels show no other indication of reaction or overgrowth with the groundmass basalt. Aluminous spinels occur as single separate single crystals, or, less commonly, as inclusions in clinopyroxene. The spinels are very

poor in Cr ($\text{Cr\#} < 1$) and have $\text{Fe}^{+2}/(\text{Fe}^{+2} + \text{Mg})$ ranging from 0.26 to 0.51. Ti and ferric iron increase with total iron. Mg# increases and ferric iron decreases with increasing Al (Figs. 6 and 7). Four individual xenocrysts were analyzed for this study and each has a distinct chemistry, with significant differences in Ti, Fe^{+2} , Fe^{+3} , Mg, and Al content (Table 3). The spinel included in clinopyroxene is distinctly higher in Cr than the other aluminous spinels (Fig. 7). A single, very small, opaque, Cr-rich spinel ($\text{Cr}/(\text{Cr} + \text{Al}) = 0.42$) was found as an inclusion in olivine (Fig. 6;

Table 3: Spinel

VMNH#	81564	81554	81554	81554	81554	81564	81564	81554	81554	81554
sample	Mh11	Mh1-5-2	Mh1-5-2	Mh1-5-1	Mh1-5-1	Mh11	Mh11	Mh1-5-2	Mh1-5-2	Mh1-5-1
n	13	33	1	23	7	3	6	1	5	4
x'l type	xenocryst	xenocryst	inclusion	xenocryst	inclusion	phenocryst	grdmass	pheno?	grdmass	grdmass
notes			in olivine		in cpx	poikilitic				
SiO ₂	0.11	0.13	0.13	0.10	0.14	0.10	0.16	0.09	0.09	0.20
Al ₂ O ₃	54.58	65.08	30.73	58.81	63.07	6.31	4.28	6.30	2.08	2.21
MgO	13.27	20.04	14.62	15.55	19.50	4.62	3.40	4.30	1.76	2.11
TiO ₂	0.96	0.27	0.62	0.55	0.31	14.65	18.75	13.95	18.09	17.78
MnO	0.20	0.10	0.13	0.17	0.10	0.53	0.66	0.51	0.85	0.83
Cr ₂ O ₃	0.14	0.13	33.51	0.12	0.62	5.43	0.45	8.05	0.13	0.12
Fe ₂ O ₃	12.18	4.38	5.48	8.06	5.09	29.83	28.59	28.71	31.39	30.80
FeO	22.15	12.87	15.19	18.47	13.06	38.25	43.47	38.30	44.40	43.33
sum	103.58	102.99	100.41	101.82	101.89	99.72	99.77	100.21	98.79	97.39
3 Cations/4 oxygens										
Si	0.003	0.003	0.004	0.003	0.004	0.003	0.006	0.003	0.003	0.008
Al	1.709	1.900	1.065	1.812	1.872	0.264	0.182	0.262	0.092	0.098
Mg	0.526	0.740	0.641	0.606	0.732	0.244	0.183	0.227	0.098	0.119
Ti	0.019	0.005	0.014	0.011	0.006	0.391	0.509	0.371	0.508	0.504
Fe ⁺³	0.243	0.082	0.121	0.159	0.097	0.796	0.776	0.764	0.882	0.874
Fe ⁺²	0.492	0.267	0.374	0.404	0.276	1.134	1.311	1.132	1.387	1.367
Mn	0.004	0.002	0.003	0.004	0.002	0.016	0.020	0.015	0.027	0.026
Cr	0.003	0.003	0.779	0.002	0.012	0.152	0.013	0.225	0.004	0.004
total	3.000	3.000	3.000	3.000	3.000	3.000	3.000	3.000	3.000	3.000
Cr/(Cr+Al)	0.002	0.001	0.422	0.001	0.007	0.351	0.064	0.462	0.039	0.036
Mg#	51.6	73.5	63.2	60.0	72.6	17.7	12.2	16.7	6.6	8.0
end-members										
Mt	0.122	0.041		0.079	0.048	0.276	0.288	0.278	0.369	0.356
Uv	0.019	0.005	0.014	0.011	0.006	0.391	0.509	0.371	0.508	0.504
Cr	0.001	0.001	0.389	0.001	0.006	0.076	0.006	0.113	0.002	0.002
Sp	0.520	0.733	0.529	0.601	0.725	0.116	0.071	0.116	0.019	0.023
Gx	0.004	0.002	0.003	0.004	0.002	0.016	0.020	0.015	0.027	0.026
Hc	0.331	0.215	0.000	0.301	0.209	0.000	0.000	0.000	0.000	0.000
Mf	0.000	0.000	0.104	0.000	0.000	0.122	0.100	0.104	0.072	0.081

Table 3). This Cr-spinel and all of the aluminous spinels plot in the fields of mantle xenoliths from alkali basalts as defined by Barnes and Roeder (2001) (Fig. 6).

Fe-Ti rich opaque spinels occur in the groundmass and as microphenocrysts in the Mole Hill microbasalt. Groundmass spinels can be classified as titano-magnetite (i.e. uv around 50 and

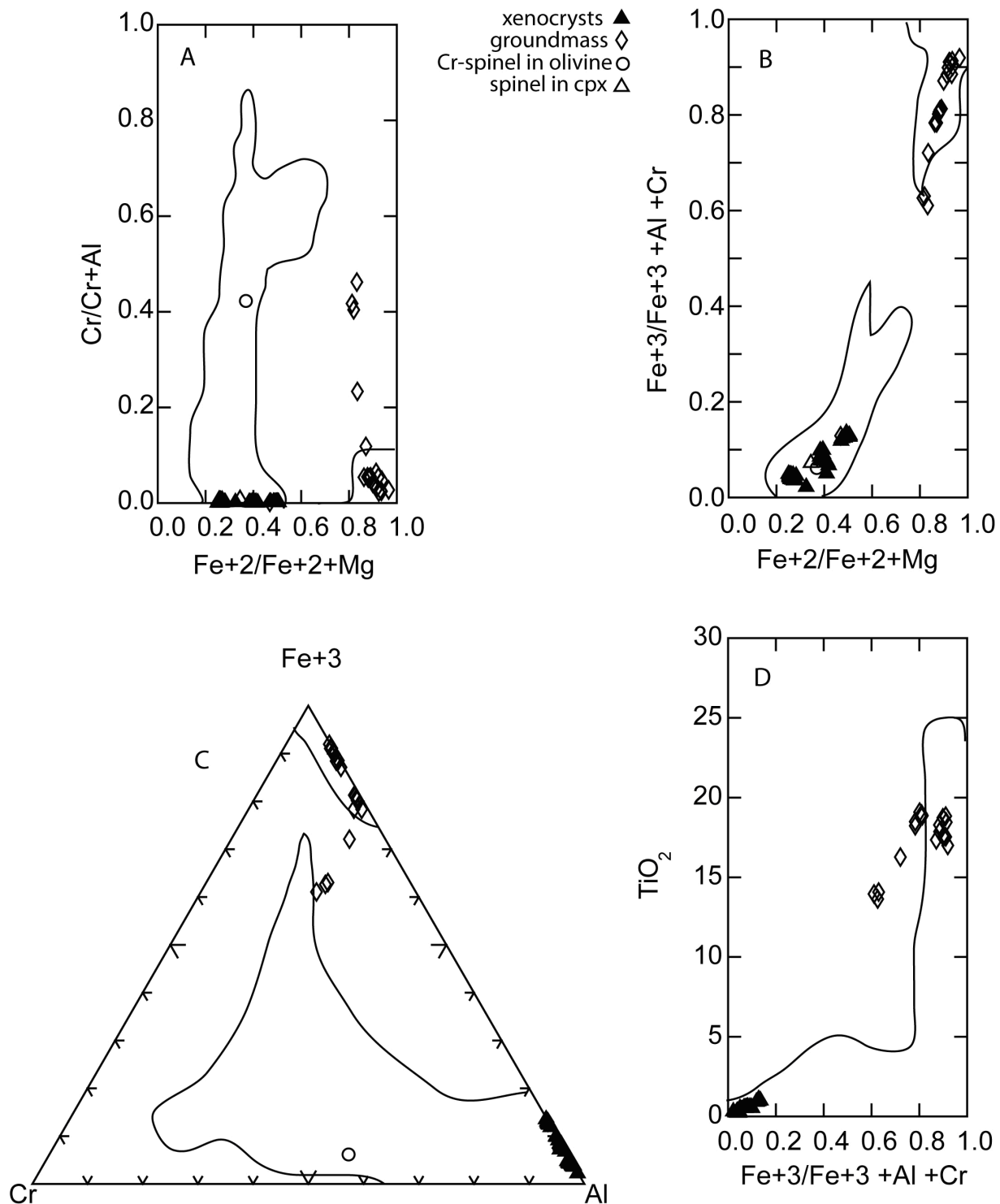


Figure 6. Spinel compositions in the Mole Hill microbasalt. Outlined fields are the fields of xenoliths in alkali basalts taken from Barnes and Roeder (2001).

mt+uv 78-92). Most have calculated magnesioferrite between 5 and 10% (Table 3). Poikilitic, opaque spinel microphenocrysts have chromite- (4-11%) and Mg-spinel- (10-12%) rich cores grading out to

titanomagnetite rims. The opaque spinels define trends of decreasing Mg and Cr and increasing followed by decreasing Ti as the magnetite component increases (Fig. 6). These spinels, at

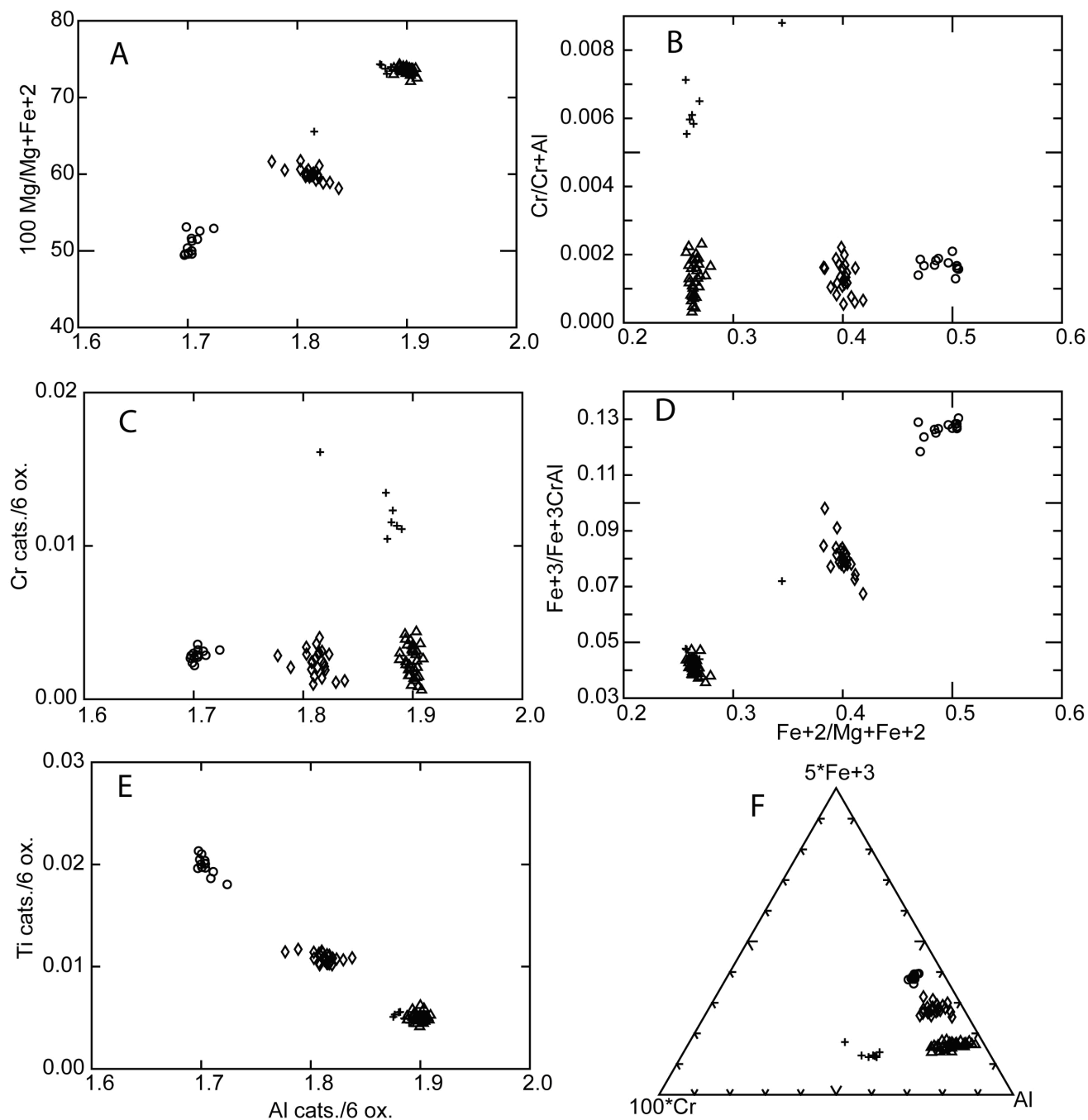


Figure 7. Aluminous Spinel xenocrysts from the Mole Hill microbasalt. Individual groups of spots represent spot analyses from single crystals. Note the elevated Cr in the spinel included in clinopyroxene.

least in part, lie outside the basalt xenolith fields as defined by Barnes and Roeder (2001).

Plagioclase. Plagioclase occurs as micro-

phenocrysts, as inclusions in the rims of sieve-textured clinopyroxene rims, and as a groundmass phase in the Mole Hill microbasalts. Plagioclase is

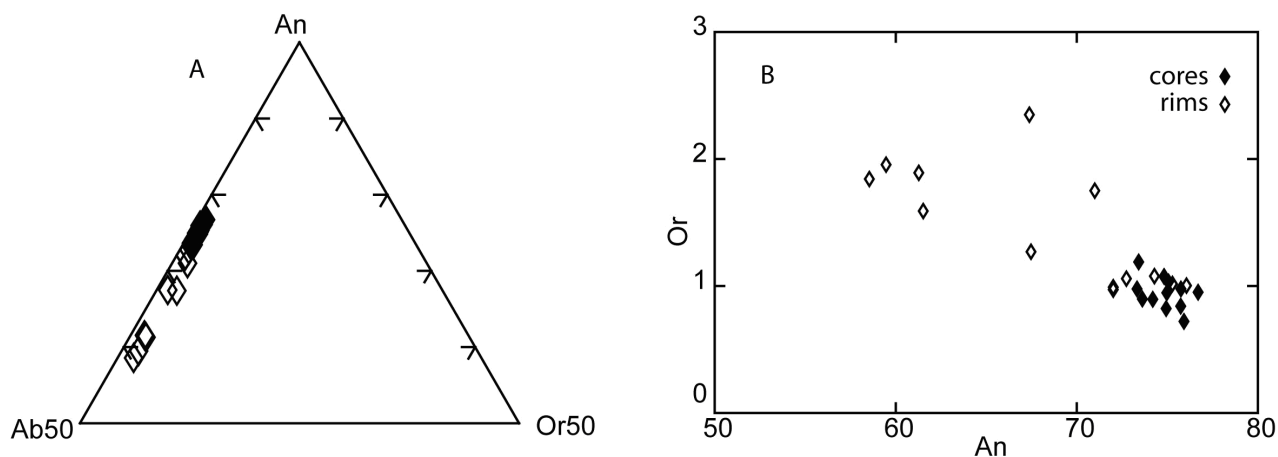


Figure 8. Plagioclase microphenocryst compositions, Mole Hill microbasalt.

Table 4: Plagioclase microphenocrysts

VMNH#	81564	81564	81554	81554	81554	81554
sample	Mh 11	Mh 11	Mh 1-5-2	Mh 1-5-2	Mh 1-5-1	Mh 1-5-1
n		11	1	8	3	10
notes	core	rim	core	rim	core	rim
SiO ₂	48.25	52.67	48.79	51.22	48.43	51.30
Al ₂ O ₃	32.17	29.71	32.23	30.76	31.85	30.15
Na ₂ O	2.84	4.71	2.88	3.83	2.88	4.07
K ₂ O	0.18	0.33	0.17	0.25	0.18	0.29
CaO	16.02	12.61	15.78	13.99	15.50	13.38
FeO	0.77	0.79	0.80	1.03	0.84	0.82
sum	100.23	100.82	100.65	101.09	99.68	100.01
Cations/8 oxygens						
Si	2.218	2.383	2.230	2.321	2.235	2.344
Al	1.743	1.584	1.736	1.643	1.732	1.624
Na	0.253	0.414	0.255	0.336	0.257	0.361
K	0.011	0.019	0.010	0.015	0.010	0.017
Ca	0.789	0.611	0.773	0.680	0.766	0.655
Fe	0.030	0.030	0.031	0.039	0.033	0.031
total	5.043	5.041	5.034	5.033	5.033	5.032
An	75.0	58.5	74.5	65.9	74.1	63.5
Ab	24.0	39.6	24.6	32.7	24.9	34.9
Or	1.0	1.8	0.9	1.4	1.0	1.6

DISCUSSION

zoned, with cores of An_{75} and narrow rims ranging down to An_{58} (Table 4, Fig. 8). Plagioclase is not observed to be in a reaction relationship with any megacryst phase.

Origin of the megacrysts and xenocrysts.

Taken as a group, the clinopyroxene and olivine megacrysts or xenocrysts from Mole Hill have compositions less evolved than typical mantle peridotites, even relatively fertile subcontinental lherzolites (e.g. Rudnick et al., 2004; Ackerman et al., 2007; Bjerg et al., 2009). On the other hand, many continental xenolith suites include a variety of clinopyroxene-rich ultramafic rocks such as clinopyroxenite, olivine clinopyroxenite, and wehrlite (Wilshire and Shervais, 1975; Irving, 1980; Ghent et al., 1980; Brearly et al., 1984; Kovacs et al., 2004; Rehfeldt et al., 2007; Xiao et al., 2010). Two groups of cpx-rich xenoliths are recognized, a Cr-diopside suite, which is typically dominated by rocks relatively rich in olivine (e.g. cpx-rich lherzolite) and an Al-augite suite in which cpx is generally the dominant mineral (Wilshire and Shervais, 1975; Irving, 1980). The Al-augite suite rocks occur as individual xenoliths, and, perhaps even more commonly, as veins intrusive into typical mantle peridotites in composite xenoliths. These cpx-rich veins reflect metasomatism of the lithosphere either by melts or fluids (Wilshire and Shervais, 1975; Irving, 1980)).

The megacrystic/xenocrystic olivine, cpx, and spinel at Mole Hill suggest an affinity with the Al-augite xenolith suite. First, the Mg# of both olivine and cpx are lower than would be expected for unmodified mantle and within the range observed in the Al-augite suite. Second, the low Cr_2O_3 (most <0.4 wt.%) and high Al_2O_3 (most >7 wt.%) in Mole Hill cpx are typical (and definitive) of the Al-augite suite (Wilshire and Shervais, 1975). Third, the Mole Hill

aluminous megacrystic spinels have exceptionally low Cr content and Cr#, among the lowest ever reported for putative mantle-derived spinels. Low Cr# (e.g. <10) spinels occur in a variety of xenolith suites (Fig. 5) (Barnes and Roeder, 2001), including Al-augite xenoliths. However, spinels with $Cr\# < 3$ appear to occur exclusively within the Al-augite suite (Wilshire and Shervais, 1975; Ghent et al., 1980; Brearly et al., 1984; Kovacs et al., 2004; Rehfeldt et al., 2007). The one feature of the Mole Hill megacryst suite that is atypical of the Al-augite suite is the relatively high NiO concentration in some Mole Hill olivine. One possible explanation for this is that the Ni-rich olivines represent fragments of the host peridotite, rather than the cpx-rich vein assemblage. The presence of a single Cr-rich spinel inclusion in one magnesian olivine megacryst is supportive of an origin distinct from the cpx and aluminous spinel megacrysts for at least some olivine.

Implications of cognate mineral compositions for origin and evolution of the picrobasalt.

At present, no reliable whole-rock chemical data exists for the Mole Hill diatreme. The single analysis in Southworth et al. (1993) is of highly altered and contaminated rock. Some hints as to the nature of the magma, however, may be found in the compositions of microphenocryst and groundmass (cognate) minerals. The moderate Cr-content of some cognate cpx and titanomagnetite, the relatively calcic nature of the plagioclase, and the moderate (e.g. ~70) Mg# of cognate cpx and olivine all suggest derivation from a somewhat, but not extensively, evolved alkali basalt-type magma. A more detailed understanding of the petrogenesis of the magma awaits further mineral and, especially, whole-rock chemistry.

ACKNOWLEDGEMENTS

Thanks to Richard Hoffman and Elizabeth Johnson for a timely and constructive reviews. The

VMNH Foundation provided partial support for this project.

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